

# PHOT 504: Quantum Photonics

## Final exam: questions & solutions

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### Exam questions

**Grading:** The final exam counts for 60% of your total grade.

**Exam type:** Closed-book, all questions can be answered **using only pen and paper**. Calculators, mobile phones, etc. are not allowed to be used during the exam.

**The duration** of the final exam is 3 hours.

This document contains both the problems and their solutions.

### Question 1: Hydrogen atom and angular momentum

Consider a **hydrogen atom** in the following superposition state  $\psi = \frac{1}{\sqrt{2}}(\psi_{210} + \psi_{211})$ .

- (a) What are the possible values (eigenvalues of  $\hat{L}_z$ ) and corresponding probabilities when measuring the z-component of the angular momentum  $L_z$ ?
- (b) Calculate the expectation value for the lowering operator  $\langle \hat{L}_- \rangle$ .

#### Solution (Q1)

(a) The hydrogen eigenstates are also eigenstates of  $L_z$  and its eigenvalues are given by  $\hat{L}_z \psi_{nlm} = m\hbar \psi_{nlm}$ :

$$\begin{aligned}\hat{L}_z \psi_{210} &= 0\hbar \psi_{210} \\ \hat{L}_z \psi_{211} &= 1\hbar \psi_{211}\end{aligned}$$

Since  $\hat{L}_z$  is Hermitian its eigenvalues are real and measurable: the eigenvalues are 0 and  $\hbar$ , and the probability to get either is the same  $|c_{210}|^2 = |c_{211}|^2 = \frac{1}{2}$ .

(b) The lowering operator of  $\hat{L}_-$  acts on the angular part of  $\psi_{nlm}$ :

$$\hat{L}_- \psi_{nlm} = \hbar \sqrt{l(l+1) - m(m-1)} \psi_{nlm}$$

The expectation value  $\langle \hat{L}_- \rangle$  is then given by:

$$\begin{aligned}
\langle \hat{L}_- \rangle &= \langle \psi | \hat{L}_- | \psi \rangle \\
&= \frac{1}{2} (\langle 2, 1, 0 | + \langle 2, 1, 1 |) \hat{L}_- (|2, 1, 0\rangle + |2, 1, 1\rangle) \\
&= \frac{1}{2} (\langle 2, 1, 0 | \hat{L}_- |2, 1, 0\rangle + \langle 2, 1, 1 | \hat{L}_- |2, 1, 0\rangle + \langle 2, 1, 0 | \hat{L}_- |2, 1, 1\rangle + \langle 2, 1, 1 | \hat{L}_- |2, 1, 1\rangle) \\
&= \frac{1}{2} (\dots \langle 2, 1, 0 | 2, 1, -1\rangle + \dots \langle 2, 1, 1 | 2, 1, -1\rangle + \hbar\sqrt{2} \langle 2, 1, 0 | 2, 1, 0\rangle + \dots \langle 2, 1, 1 | 2, 1, 0\rangle) \\
&= \frac{1}{2} (0 + 0 + \hbar\sqrt{2} \langle 2, 1, 0 | 2, 1, 0\rangle + 0) \\
&= \frac{\hbar}{\sqrt{2}}
\end{aligned}$$

Where the third term is the only nonzero one due to orthonormality of the eigenstates.

## Question 2: Spin in a magnetic field

Consider a spin 1/2 particle in a magnetic field oriented along the z-axis:  $\vec{B} = B\vec{e}_z$ , has following time-independent Pauli (Schrodinger) equation with eigenstates  $\chi_1, \chi_2$  and eigenenergies  $E_1, E_2$  are given by:

$$\mu_B B \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \begin{pmatrix} u \\ d \end{pmatrix} = E \begin{pmatrix} u \\ d \end{pmatrix} \quad \begin{cases} \text{spin-up:} & E_1 = \mu_B B, & \chi_1 = \begin{pmatrix} 1 \\ 0 \end{pmatrix} \\ \text{spin-down:} & E_2 = -\mu_B B, & \chi_2 = \begin{pmatrix} 0 \\ 1 \end{pmatrix} \end{cases}$$

where  $\mu_B = \frac{e\hbar}{2m_0}$  is the Bohr magneton. Assume at time zero  $t = 0$  the system is in the state:  $\chi(0) = \frac{1}{\sqrt{5}} (2\chi_1 + i\chi_2) = \frac{1}{\sqrt{5}} \begin{pmatrix} 2 \\ i \end{pmatrix}$ .

(a) Derive the time-dependent probability density function:  $|\chi(t)|^2$ , and simplify.

(b) Write down the expression for the time-dependent expectation value  $\langle \hat{S}_x \rangle = \langle \chi(t) | \hat{S}_x | \chi(t) \rangle$  and simplify.

## Solution (Q2)

(a) The time-dependent probability density function  $|\chi(t)|^2$  can be calculated by adding the time-dependent factors  $e^{-iE_n t/\hbar}$  of the eigenstates. To simplify we denote  $\omega = E_1/\hbar = -E_2/\hbar$ , the time-dependent spinor becomes then:

$$\begin{aligned}
\chi(t) &= \begin{pmatrix} u(t) \\ d(t) \end{pmatrix} = \frac{1}{\sqrt{5}} (2\chi_1 e^{-iE_1 t/\hbar} + i\chi_2 e^{-iE_2 t/\hbar}) \\
&= \frac{1}{\sqrt{5}} \left( 2e^{-i\omega t} \begin{pmatrix} 1 \\ 0 \end{pmatrix} + ie^{i\omega t} \begin{pmatrix} 0 \\ 1 \end{pmatrix} \right) \\
&= \frac{1}{\sqrt{5}} \begin{pmatrix} 2e^{-i\omega t} \\ ie^{i\omega t} \end{pmatrix}
\end{aligned}$$

and thus the probability density functions of the individual spin-components  $u$  and  $d$  of  $\chi = \begin{pmatrix} u \\ d \end{pmatrix}$  are independent of time and given by:

$$|u(t)|^2 = \frac{1}{5} |2e^{-i\omega t}|^2 = \frac{4}{5}, \quad |d(t)|^2 = \frac{1}{5} |ie^{i\omega t}|^2 = \frac{1}{5}$$

And the total probability is normalized as should:  $|u|^2 + |d|^2 = 1$ .

(b) For the time-dependent expectation value  $\langle \hat{S}_x \rangle = \langle \chi(t) | \hat{S}_x | \chi(t) \rangle$  we use the matrix representation for  $\chi(t) = \frac{1}{\sqrt{5}} \begin{pmatrix} 2e^{-i\omega t} \\ ie^{i\omega t} \end{pmatrix}$  calculated in (a), and  $\hat{S}_x = \frac{\hbar}{2} \sigma_x = \frac{\hbar}{2} \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}$ :

$$\begin{aligned}
\langle \chi(t) | \hat{S}_x | \chi(t) \rangle &= (u^*(t) \quad d^*(t)) \frac{\hbar}{2} \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \begin{pmatrix} u(t) \\ d(t) \end{pmatrix} \\
&= \frac{\hbar}{2} \frac{1}{5} (2e^{i\omega t} \quad -ie^{-i\omega t}) \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \begin{pmatrix} 2e^{-i\omega t} \\ ie^{i\omega t} \end{pmatrix} \\
&= \frac{\hbar}{10} (2e^{i\omega t} \quad -ie^{-i\omega t}) \begin{pmatrix} ie^{i\omega t} \\ 2e^{-i\omega t} \end{pmatrix} \\
&= \frac{\hbar}{10} i2 (e^{i2\omega t} - e^{i2\omega t}) \\
&= \frac{\hbar}{10} 4i^2 \sin(2\omega t) = -\frac{2\hbar}{5} \sin(2\omega t)
\end{aligned}$$

Filling in the value for the energy  $\omega = \mu_B B / \hbar$  we get:  $\langle \hat{S}_x \rangle = -\frac{2\hbar}{5} \sin\left(\frac{2\mu_B B}{\hbar} t\right)$ .

### Question 3: Perturbation of a Three-State System

Consider a three-state system:  $\hat{H} = \hat{H}_0 + \hat{H}_p$  with unperturbed Hamiltonian  $H_0$  and perturbation term  $H_p$  in matrix representation given by:

$$H = H_0 + H_p, \quad H_0 = \begin{pmatrix} -1 & 0 & 0 \\ 0 & 0 & 2 \\ 0 & 2 & 0 \end{pmatrix}, \quad H_p = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & -1 \end{pmatrix}$$

(a) Find the eigenenergies ( $E_1^{(0)}$ ,  $E_2^{(0)}$ ,  $E_3^{(0)}$ ), and corresponding eigenstates ( $\psi_1^{(0)}$ ,  $\psi_2^{(0)}$ ,  $\psi_3^{(0)}$ ) of the unperturbed Hamiltonian by solving the eigenvalue equation  $H\psi = E\psi$ .

(b) Then calculate the perturbed energy values using first order perturbation theory:

$$E_n = E_n^{(0)} + E_n^{(1)}, \quad E_n^{(1)} = \langle \psi_n^{(0)} | H_p | \psi_n^{(0)} \rangle$$

### Solution (Q3)

To simplify the problem, the units of the Hamiltonian were chosen left (energy)

(a) The eigenenergies  $E_1^{(0)}$ ,  $E_2^{(0)}$ , and  $E_3^{(0)}$ , can be found by solving the eigenvalue equation  $H_0\psi = \lambda\psi$  with solutions  $\psi_n$  column vectors in  $\mathbb{C}^3$ .

$$(\lambda\mathbb{1} - H_0)\psi = 0 \quad \Rightarrow \quad \det(\lambda\mathbb{1} - H_0) = 0$$

We can factorize the determinant to obtain solution for  $\lambda$ :

$$\begin{vmatrix} \lambda + 1 & 0 & 0 \\ 0 & \lambda & -2 \\ 0 & -2 & \lambda \end{vmatrix} = (\lambda + 1)(\lambda^2 - 4) = (\lambda + 1)(\lambda + 2)(\lambda - 2)$$

and eigenenergies are thus  $E_1^{(0)} = -1$ ,  $E_2^{(0)} = -2$ ,  $E_3^{(0)} = 2$ .

For the eigenstates we get:

$$\boxed{E_1^{(0)} = -1}:$$

$$\begin{pmatrix} -1 & 0 & 0 \\ 0 & 0 & 2 \\ 0 & 2 & 0 \end{pmatrix} = -1 \begin{pmatrix} x \\ y \\ z \end{pmatrix} \quad \Rightarrow \quad x = x, y = z = 0 \quad \Rightarrow \quad \psi_1^{(0)} = \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix}$$

$$\boxed{E_2^{(0)} = -2}:$$

$$\begin{pmatrix} -1 & 0 & 0 \\ 0 & 0 & 2 \\ 0 & 2 & 0 \end{pmatrix} = -2 \begin{pmatrix} x \\ y \\ z \end{pmatrix} \quad \Rightarrow \quad x = 0, z = -y \quad \Rightarrow \quad \psi_2^{(0)} = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ 1 \\ -1 \end{pmatrix}$$

$$\boxed{E_3^{(0)} = 2}:$$

$$\begin{pmatrix} -1 & 0 & 0 \\ 0 & 0 & 2 \\ 0 & 2 & 0 \end{pmatrix} = 2 \begin{pmatrix} x \\ y \\ z \end{pmatrix} \quad \Rightarrow \quad x = 0, z = y \quad \Rightarrow \quad \psi_3^{(0)} = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ 1 \\ 1 \end{pmatrix}$$

(b) To find the perturbed energies up to 1th order

$$E_1^{(1)} = \langle \psi_1^{(0)} | H_p | \psi_1^{(0)} \rangle = (1 \ 0 \ 0) \begin{pmatrix} 0 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & -1 \end{pmatrix} \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix} = (1 \ 0 \ 0) \begin{pmatrix} 0 \\ 0 \\ 0 \end{pmatrix} = 0$$

$$E_2^{(1)} = \langle \psi_2^{(0)} | H_p | \psi_2^{(0)} \rangle = (0 \ 1 \ -1) \begin{pmatrix} 0 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & -1 \end{pmatrix} \begin{pmatrix} 0 \\ 1 \\ -1 \end{pmatrix} = (0 \ 1 \ -1) \begin{pmatrix} 0 \\ 1 \\ 1 \end{pmatrix} = 0$$

$$E_3^{(1)} = \langle \psi_3^{(0)} | H_p | \psi_3^{(0)} \rangle = (0 \ 1 \ 1) \begin{pmatrix} 0 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & -1 \end{pmatrix} \begin{pmatrix} 0 \\ 1 \\ 1 \end{pmatrix} = (0 \ 1 \ 1) \begin{pmatrix} 0 \\ 1 \\ -1 \end{pmatrix} = 0$$

Up to first order the eigenenergies do not differ in this case:  $E_1 = E_1^{(0)} = -1$ ,  $E_2 = E_2^{(0)} = -2$ ,  $E_3 = E_3^{(0)} = 2$ .

The actual energies (**not required for this question**) can be obtained from directly solving the perturbed Hamiltonian and are different:

$$\begin{vmatrix} \lambda + 1 & 0 & 0 \\ 0 & \lambda - 1 & -2 \\ 0 & -2 & \lambda + 1 \end{vmatrix} = (\lambda + 1) [(\lambda^2 - 1) - 4] = (\lambda + 1)(\lambda^2 - 5)$$

and the actual eigenenergies are thus  $E_1 = -1$ ,  $E_2 = -\sqrt{5}$ ,  $E_3 = \sqrt{5}$ .

#### Question 4: Time-reversal symmetry of a spin 1/2 particle

Consider a spin 1/2 particle defined in the standard basis of spin-up and spin-down:  $\{|\uparrow\rangle, |\downarrow\rangle\}$ . The time-reversal operator  $\Theta$  acting on a spinor has the following effect:

$$\Theta \begin{pmatrix} u \\ d \end{pmatrix} = \begin{pmatrix} -d^* \\ u^* \end{pmatrix}$$

(a) Assume the particle is in the spin-up state:  $\chi_u = \begin{pmatrix} 1 \\ 0 \end{pmatrix}$ , in what state is the particle after you apply the time-reversal operator  $\Theta$ ?

(b) Suppose now the particle is in state  $\chi = \begin{pmatrix} 1 \\ 2i \end{pmatrix}$ . What is the expectation value  $\langle \Theta \rangle = \langle \chi | \Theta | \chi \rangle$ ?

#### Solution (Q4)

(a) Applying the time-reversal operator to the spin-up state results in the spin-down state:

$$\Theta \chi_u = \Theta \begin{pmatrix} 1 \\ 0 \end{pmatrix} = \begin{pmatrix} 0 \\ 1 \end{pmatrix}$$

(b) We can calculate the expectation value of the time-reversal operator for a more general spinor  $\chi = \begin{pmatrix} u \\ d \end{pmatrix}$ :

$$\langle \Theta \rangle = \langle \chi | \Theta | \chi \rangle = (u^* \quad d^*) \Theta \begin{pmatrix} u \\ d \end{pmatrix} = (u^* \quad d^*) \begin{pmatrix} -d^* \\ u^* \end{pmatrix} = -u^* d^* + d^* u^* = 0$$

This result is also valid for  $\chi = \begin{pmatrix} 1 \\ 2i \end{pmatrix}$  and further tells us that  $\Theta|\chi\rangle$  and  $|\chi\rangle$  are orthogonal. btw: this is true for general spinors described by the Pauli-Schrodinger equation.

## Question 5: Life time of a particle in a hydrogen atom

The life-time of a particle is given by  $\tau = 1/A$ . Assume a particle is in the  $p_z$  orbital, i.e.  $\psi_{2,1,0}$  (ignore spin) and consider the life-time for falling back to the ground state  $\psi_{1,0,0}$ . In this case  $\mathcal{P}_x$  and  $\mathcal{P}_y$  are zero and  $|\mathcal{P}|^2 = |\mathcal{P}_z|^2$ . In this case  $A$  is given by:

$$A = \frac{\Delta\omega^3}{3\pi\epsilon_0\hbar c^3} |\mathcal{P}_z|^2$$

- (a) First prove that  $\mathcal{P}_x = q\langle\psi_{1,0,0}|x|\psi_{2,1,0}\rangle = 0$ .  
 (b) Then calculate  $|\mathcal{P}_z|^2 = q^2|\langle\psi_{1,0,0}|z|\psi_{2,1,0}\rangle|^2$ .

### Solution (Q5)

The eigenstates are given by (both cartesian and spherical coordinates):

$$\begin{aligned}\psi_{100}(r, \theta, \phi) &= \frac{1}{\sqrt{\pi a^3}} e^{-r/a}, & \psi_{100}(x, y, z) &= \frac{1}{\sqrt{\pi a^3}} e^{-\sqrt{x^2+y^2+z^2}/a} \\ \psi_{210}(r, \theta, \phi) &= \frac{1}{4\sqrt{2\pi a^3}} \frac{r}{a} e^{-r/2a} \cos\theta, & \psi_{210}(x, y, z) &= \frac{1}{4\sqrt{2\pi a^3}} \frac{z}{a} e^{-\sqrt{x^2+y^2+z^2}/2a}\end{aligned}$$

(a) To prove that  $\mathcal{P}_x = q\langle\psi_{1,0,0}|x|\psi_{2,1,0}\rangle = 0$  we can argue using symmetry around the z-axis of both  $|\psi_{2,1,0}\rangle$  and  $|\psi_{1,0,0}\rangle$  and the integral over the volume with positive  $x$  will compensate for the volume with negative  $x$ :

$$\begin{aligned}\langle\psi_{1,0,0}|x|\psi_{2,1,0}\rangle &= \iiint_{\mathbb{R}^3} x\psi_{100}^*\psi_{210} dx dy dz \\ &= \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} \left( \int_{-\infty}^0 x\psi_{100}^*\psi_{210} dx + \int_0^{+\infty} x\psi_{100}^*\psi_{210} dx \right) dy dz \\ &= \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} \left( \int_0^{+\infty} (-x)\psi_{100}^*(-x, y, z)\psi_{210}(-x, y, z) dx \right. \\ &\quad \left. + \int_0^{+\infty} x\psi_{100}^*(x, y, z)\psi_{210}(x, y, z) dx \right) dy dz \\ &= \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} \left( - \int_0^{+\infty} x\psi_{100}^*(x, y, z)\psi_{210}(x, y, z) dx \right. \\ &\quad \left. + \int_0^{+\infty} x\psi_{100}^*(x, y, z)\psi_{210}(x, y, z) dx \right) dy dz \\ &= \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} (0) dy dz = 0\end{aligned}$$

Where we used the fact that  $\psi_{100}$  and  $\psi_{210}$  are real, and symmetric in  $x$  around  $x = 0$ :

$$\begin{aligned}\psi_{100}^*(-x, y, z) &= \psi_{100}^*(x, y, z) = \psi_{100}(x, y, z) \\ \psi_{210}(-x, y, z) &= \psi_{210}(x, y, z)\end{aligned}$$

*Alternative solution:* We can get to the same result by using spherical coordinates and  $x = r \sin \theta \cos \phi$ :

$$\begin{aligned}\langle \psi_{1,0,0} | x | \psi_{2,1,0} \rangle &= \iiint_{\mathbb{R}^3} r \sin \theta \cos \phi \psi_{100}^* \psi_{210} r^2 \sin \theta dr d\theta d\phi \\ &= \int_0^{+\infty} r^3 \frac{1}{\sqrt{\pi a^3}} e^{-r/a} \frac{1}{4\sqrt{2\pi a^3}} \frac{r}{a} e^{-r/2a} dr \int_0^\pi \cos \theta \sin^2 \theta d\theta \int_0^{2\pi} \cos \phi d\phi \\ &= \frac{1}{4\sqrt{2\pi a^4}} \int_0^{+\infty} r^4 e^{-3r/2a} dr \int_0^\pi \cos \theta \sin^2 \theta d\theta \int_0^{2\pi} \cos \phi d\phi = 0\end{aligned}$$

Since the last factor is zero (and the other factors are finite) the whole integral becomes zero:

$$\int_0^{2\pi} \cos \phi d\phi = [\sin \phi] \Big|_0^{2\pi} = 0$$

(b) To calculate the z-component of  $\mathcal{P}$ :  $\mathcal{P}_z = q \langle \psi_{1,0,0} | z | \psi_{2,1,0} \rangle$  we use integration in spherical coordinates and use  $z = r \cos \theta$ :

$$\begin{aligned}\langle \psi_{1,0,0} | z | \psi_{2,1,0} \rangle &= \iiint_{\mathbb{R}^3} r \psi_{100}^* \psi_{210} r^2 \sin \theta \cos \theta dr d\theta d\phi \\ &= \int_0^{+\infty} r^3 \frac{1}{\sqrt{\pi a^3}} e^{-r/a} \frac{1}{4\sqrt{2\pi a^3}} \frac{r}{a} e^{-r/2a} dr \cdot \int_0^\pi \cos^2 \theta \sin \theta d\theta \cdot \int_0^{2\pi} d\phi \\ &= \frac{1}{4\sqrt{2\pi a^4}} \int_0^{+\infty} r^4 e^{-3r/2a} dr \cdot \left[ -\frac{1}{2+1} \cos^3 \theta \right] \Big|_0^\pi \cdot 2\pi \\ &= \frac{1}{4\sqrt{2\pi a^4}} \frac{4!}{(3/2a)^5} \cdot \frac{2}{3} \cdot 2\pi \\ &= \frac{2^8}{3^5 \sqrt{2}} \cdot a\end{aligned}$$

Then we take the absolute value square:  $|\mathcal{P}_z|^2 = \frac{2^{15}}{3^{10}} \cdot q^2 a^2$ .

**This last part is not required for this question:** This value could then be used to calculate the spontaneous emission rate  $A$  where we fill in  $\Delta\omega = -\frac{3}{4} \frac{\text{Ry}}{\hbar}$  and apply once  $\text{Ry} = \frac{me^4}{8\epsilon_0^2 \hbar^2}$  and fill in the Bohr radius  $a = \frac{4\pi\epsilon_0 \hbar^2}{me^2}$  again once:

$$\begin{aligned}A &= \frac{\Delta\omega^3}{3\pi\epsilon_0 \hbar c^3} |\mathcal{P}_z|^2 = \frac{\Delta\omega^3}{3\pi\epsilon_0 \hbar c^3} \cdot \frac{2^{15}}{3^{10}} \cdot e^2 a^2 = \frac{\left(\frac{3}{4}\text{Ry}\right)^3}{3\pi\epsilon_0 \hbar^4 c^3} \cdot \frac{2^{15}}{3^{10}} \cdot e^2 a^2 = \frac{\text{Ry}^3}{\pi\epsilon_0 \hbar^4 c^3} \cdot \frac{2^9}{3^8} \cdot e^2 a^2 \\ &= \frac{\text{Ry}^3}{\hbar^4 c^3} \cdot \frac{2^9}{3^8} \cdot \frac{4\hbar^2}{m} a = \frac{\text{Ry}^2}{m^2 c^4} \cdot \frac{2^{10}}{3^8} \cdot \frac{c}{a}\end{aligned}$$

Remembering that  $mc^2 = 0.511 \text{ MeV}$  we obtain:

$$A = \frac{\text{Ry}^2}{m^2 c^4} \cdot \frac{2^{10}}{3^8} \cdot \frac{c}{a} = \left( \frac{13.6 \text{ eV}}{0.511 \times 10^6 \text{ eV}} \right)^2 \cdot \frac{2^{10}}{3^8} \cdot \frac{3 \times 10^8 \text{ m/s}}{0.529 \times 10^{-10} \text{ m}} \approx 6.27 \times 10^8 \text{ s}^{-1}$$

Therefore  $\tau = 1/A \approx 1.6 \times 10^{-9} \text{ s}$